



“Tell me that you have found no sign of
New Physics again, I dare you.
I double dare you. Tell me
one more goddamn **time!**”

Electroweak Precision Observables and BSM Physics

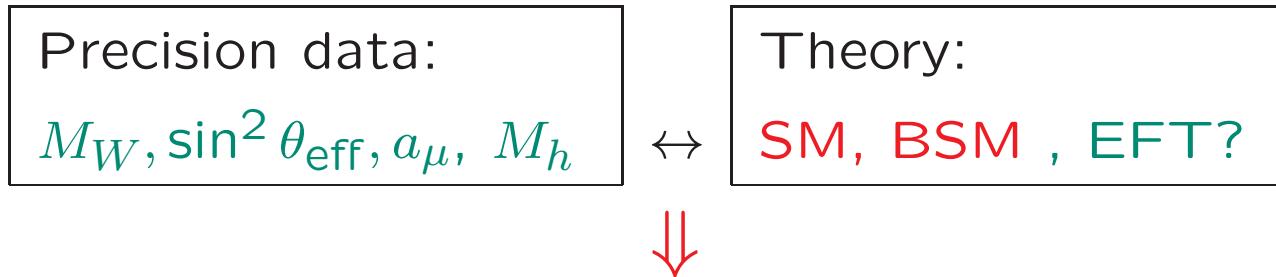
Sven Heinemeyer, IFT/IFCA (CSIC, Madrid/Santander)

virtual only, 07/2020

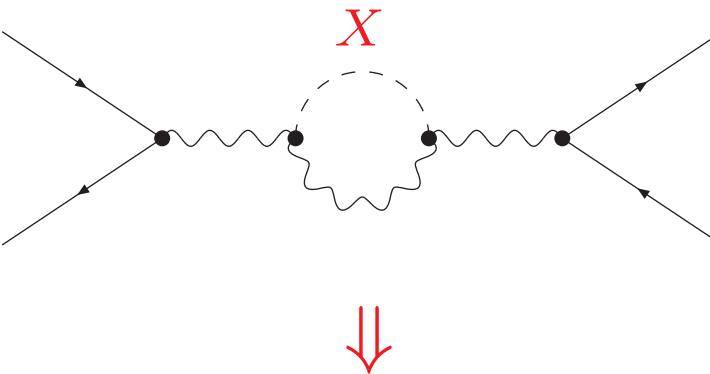
1. Introduction: Electroweak Precision Observables
2. EWPOs in concrete BSM theories
3. Why not (only) EFT?
4. Conclusions

1. Introduction: Electroweak Precision Observables

Comparison of observables with theory:



Test of theory at quantum level: Sensitivity to loop corrections, e.g. X



SM: limits on M_H , BSM: limits on M_X

Very high accuracy of measurements and theoretical predictions needed
⇒ only models “ready” so far: SM, MSSM and maybe pure Higgs extensions

The “classical” EWPO:

$$M_W \quad (\text{best from } e^+e^- \text{ threshold scan})$$

$$\sigma_{\text{had}}^0 = \sum_q \sigma_q(M_Z^2),$$

$$\Gamma_Z = \sum_f \Gamma[Z \rightarrow f\bar{f}], \quad (\text{from a fit to } \sigma_f(s) \text{ at various values of } s)$$

$$R_\ell = [\sum_q \sigma_q(M_Z^2)] / \sigma_\ell(M_Z^2), \quad (\ell = e, \mu, \tau)$$

$$R_q = \sigma_q(M_Z^2) / [\sum_q \sigma_q(M_Z^2)], \quad (q = b, c)$$

$$A_{\text{FB}}^f = \frac{\sigma_f(\theta < \frac{\pi}{2}) - \sigma_f(\theta > \frac{\pi}{2})}{\sigma_f(\theta < \frac{\pi}{2}) + \sigma_f(\theta > \frac{\pi}{2})} \equiv \frac{3}{4} \mathcal{A}_e \mathcal{A}_f,$$

$$A_{\text{LR}}^f = \frac{\sigma_f(P_e < 0) - \sigma_f(P_e > 0)}{\sigma_f(P_e < 0) + \sigma_f(P_e > 0)} \equiv \mathcal{A}_e |P_e|$$

$$\mathcal{A}_f = 2 \frac{g_{V_f}/g_{A_f}}{1 + (g_{V_f}/g_{A_f})^2} = \frac{1 - 4|Q_f| \sin^2 \theta_{\text{eff}}^f}{1 - 4|Q_f| \sin^2 \theta_{\text{eff}}^f + 8(|Q_f| \sin^2 \theta_{\text{eff}}^f)^2} \quad (f = \ell, b, \dots)$$

More “possible” EWPO:

- the anomalous magnetic moment of the muon: a_μ
- Higgs boson masses ~ 125 GeV
 - ⇒ if M_h is not a free parameter
 - Challenge:** before the end of this talk think of a non-SUSY BSM theory that predicts M_h !
- new **couplings** between SM and BSM
- ...

⇒ focus here mostly on the “classical” EWPO

Evaluation of “classical” EWPO in the SM (and BSM)

A) Theoretical prediction for M_W in terms

of $M_Z, \alpha, G_\mu, \Delta r$:

$$M_W^2 \left(1 - \frac{M_W^2}{M_Z^2} \right) = \frac{\pi \alpha}{\sqrt{2} G_\mu} (1 + \Delta r)$$

\Updownarrow
loop corrections

Evaluate Δr from μ decay $\Rightarrow M_W$

One-loop result for M_W in the SM:

[A. Sirlin '80] , [W. Marciano, A. Sirlin '80]

$$\begin{aligned} \Delta r_{\text{1-loop}} &= \Delta \alpha - \frac{c_W^2}{s_W^2} \Delta \rho + \Delta r_{\text{rem}}(M_H) \\ &\sim \log \frac{M_Z}{m_f} \quad \sim m_t^2 \quad \log(M_H/M_W) \\ &\sim 6\% \quad \sim 3.3\% \quad \sim 1\% \end{aligned}$$

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\Updownarrow
loop corrections

B) Effective mixing angle:

$$\sin^2 \theta_{\text{eff}} = \frac{1}{4 |Q_f|} \left(1 - \frac{\text{Re } g_V^f}{\text{Re } g_A^f} \right)$$

Higher order contributions:

$$g_V^f \rightarrow g_V^f + \Delta g_V^f, \quad g_A^f \rightarrow g_A^f + \Delta g_A^f$$

2. EWPOs in concrete BSM models

The by far best worked out model: **SM**

Intrinsic uncertainties:

Quantity	current experimental unc.	current intrinsic unc.
M_W [MeV]	12	4 ($\alpha^3, \alpha^2 \alpha_s$)
$\sin^2 \theta_{\text{eff}}^\ell$ [10^{-5}]	16	4.5 ($\alpha^3, \alpha^2 \alpha_s$)
Γ_Z [MeV]	2.3	0.5 ($\alpha^3, \alpha^2 \alpha_s, \alpha \alpha_s^2$)
R_b [10^{-5}]	66	15 ($\alpha^3, \alpha^2 \alpha_s$)
R_l [10^{-3}]	25	5 ($\alpha^3, \alpha^2 \alpha_s$)

Parametric uncertainties:

Quantity	$\delta m_t = 0.9$ GeV	$\delta(\Delta \alpha_{\text{had}}) = 10^{-4}$	$\delta M_Z = 2.1$ MeV
δM_W^{para} [MeV]	5.5	2	2.5
$\delta \sin^2 \theta_{\text{eff}}^{\ell, \text{para}}$ [10^{-5}]	3.0	3.6	1.4

⇒ Current intrinsic/parametric uncertainties are substantially smaller than current experimental uncertainties :-) **in the SM!**

Which BSM theories have been sufficiently worked out?

What means “sufficiently”?

- several/all EWPOs are available at full 1-loop
- at best: leading 2-loop
- uncertainty estimate available

MSSM: all EWPO at full 1-loop, $\Delta\rho$ 2-loop: $\Delta\rho^{\alpha\alpha_s}$, $\Delta\rho^{\alpha_t^2, \alpha_t\alpha_b, \alpha_b^2}$
rough uncertainty estimate available

NMSSM: only M_W , $\Delta\rho^{1\text{-loop}}$, $\Delta\rho^{\alpha\alpha_s}$, no uncertainty estimate

xSM: only S, T, U at 1-loop, no uncertainty estimate

2HDMs: only S, T, U at 1-loop, T at 2-loop, no uncertainty estimate

⇒ better overview necessary! ⇒ Snowmass 2021?

Pure Higgs sector extensions:

S, T, U probably sufficient, still uncertainty estimate needed

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- What BSM models have been worked out?
- ... and to what extent?
- Uncertainty estimate?
- What has to be calculated to have them “sufficiently” under control?
(... any volunteers? :-)
- Any possible predictions about other parts of the BSM spectrum?

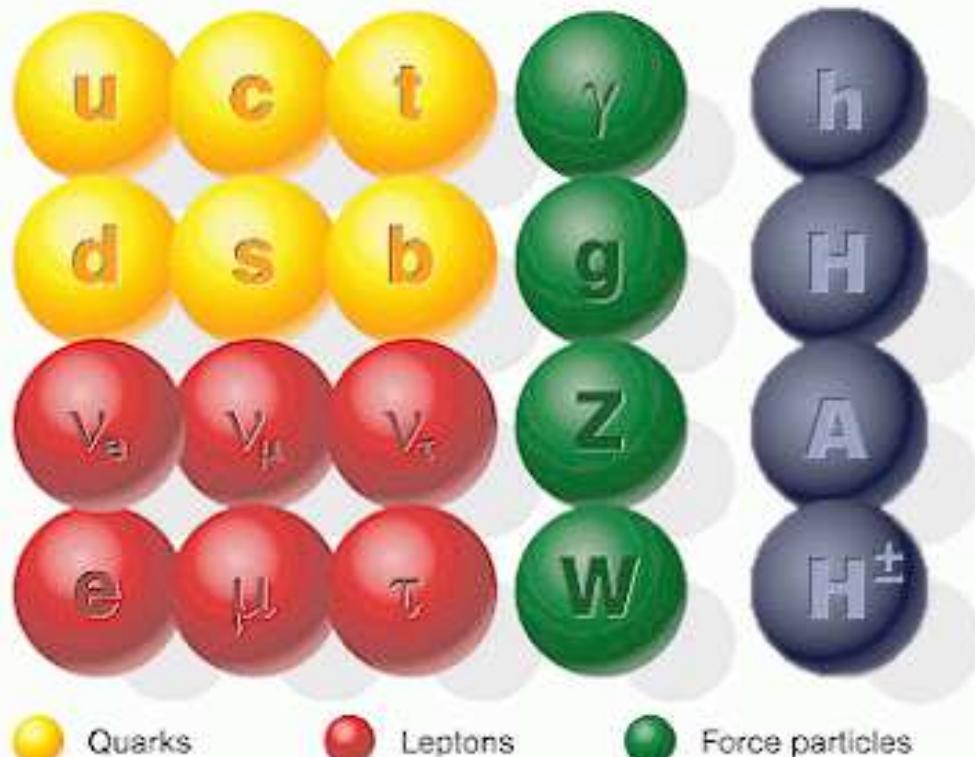
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- What has to be calculated to have them “sufficiently” under control?
(... any volunteers? :-)
- Any possible predictions about other parts of the BSM spectrum?
- Are there some clear patterns arising?
... that can be compared to future data?

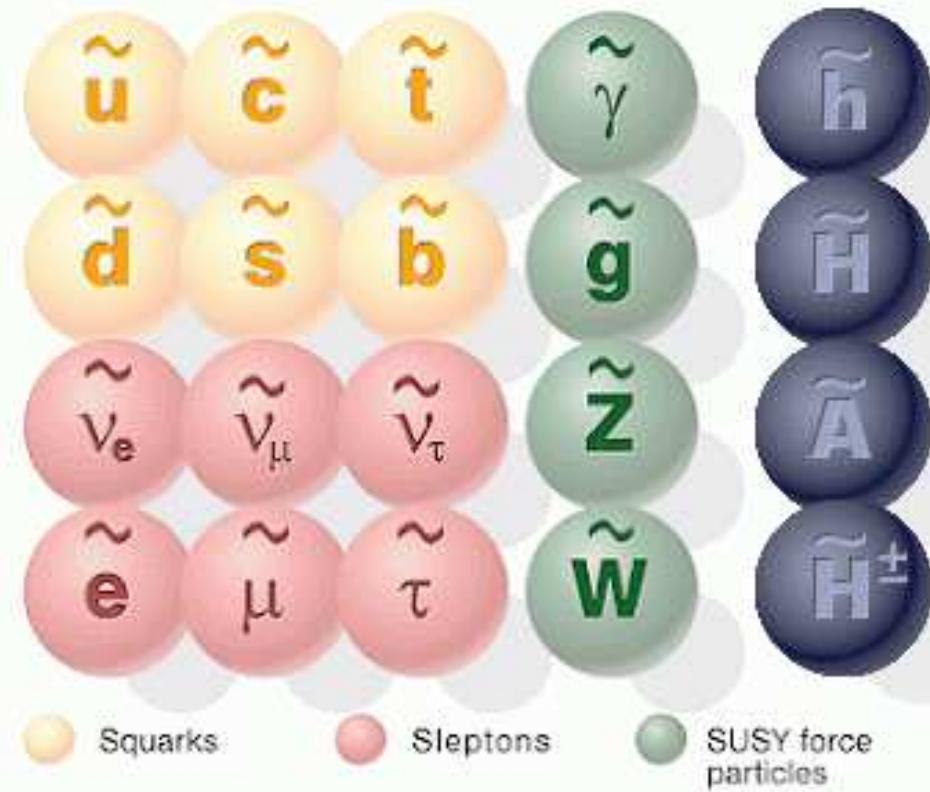
EWPOs in the MSSM

Superpartners for Standard Model particles

Standard particles



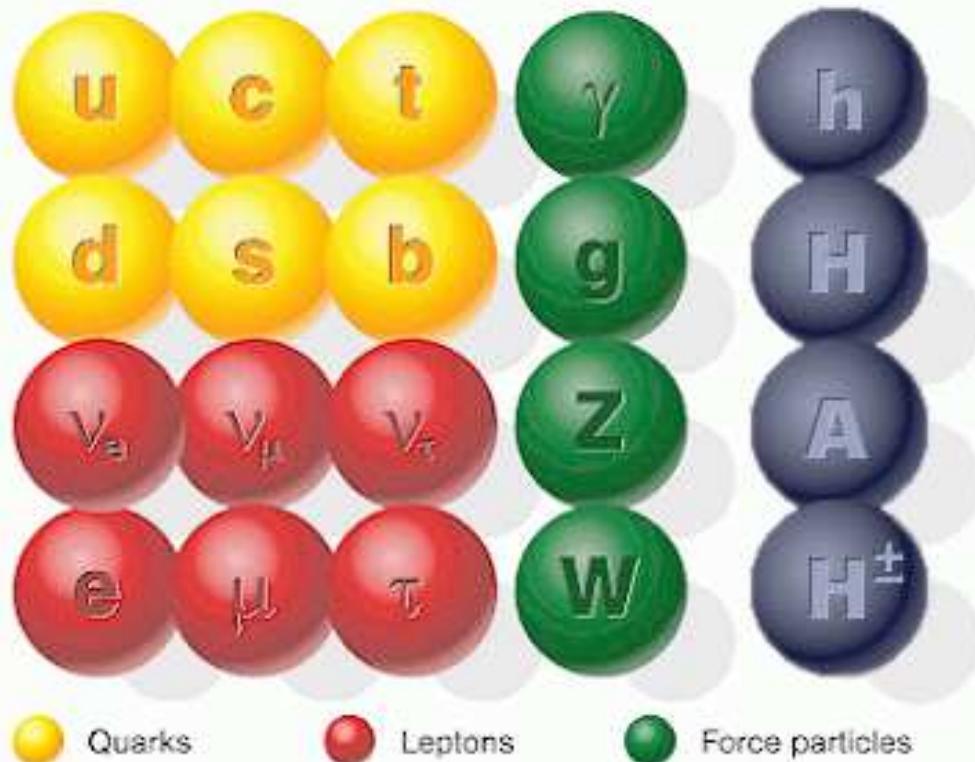
SUSY particles



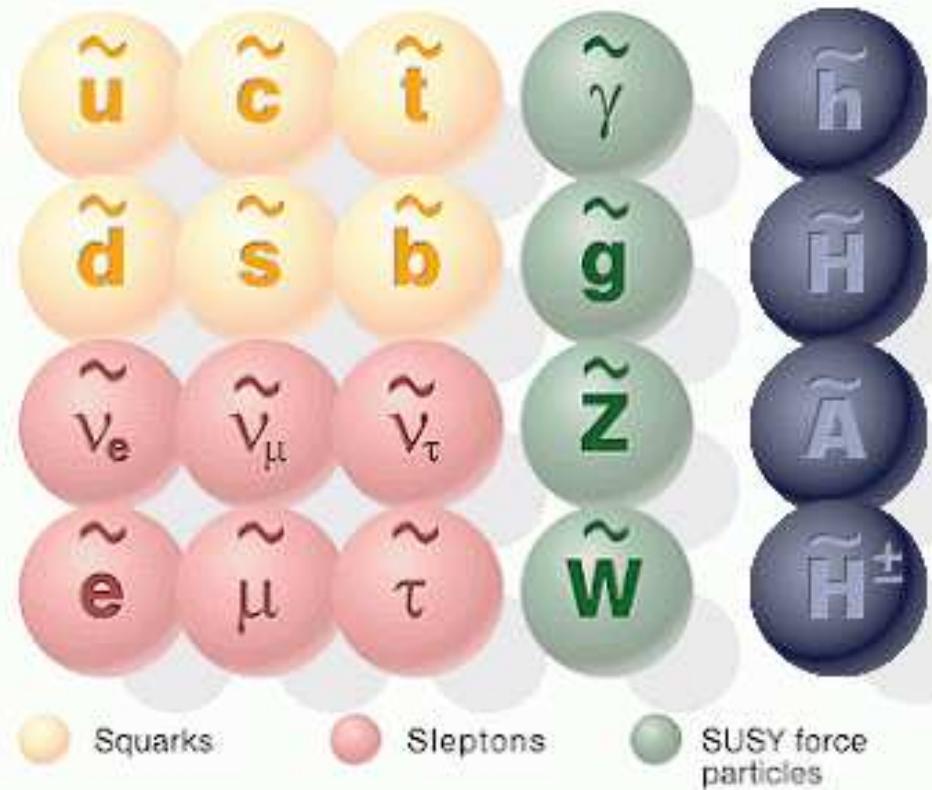
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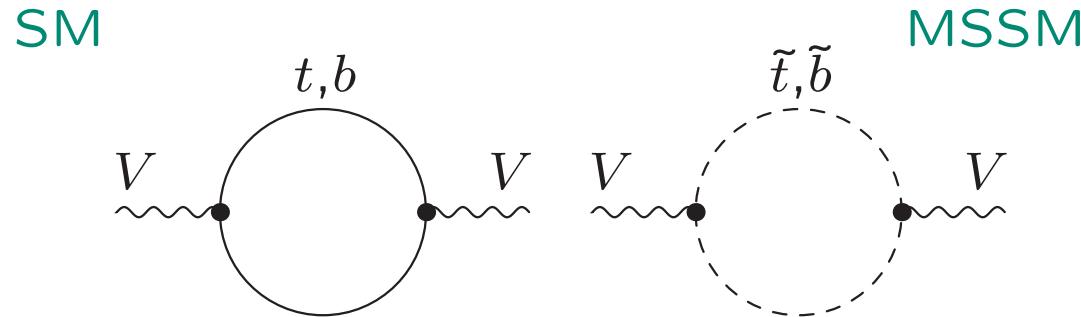
SUSY particles



Repeat the following excercises in your favorite BSM theory!

Differences between the MSSM and the SM:

1.) New contributions from SUSY particles:



2.) CPV effects via new CPV phases

3.) large Yukawa corrections: $\sim m_t^4 \log \left(\frac{m_{\tilde{t}_1} m_{\tilde{t}_2}}{m_t^2} \right)$

4.) large corrections from the b/\tilde{b} sector for large $\tan \beta$

5.) non-decoupling SUSY effects: $\sim \log \frac{M_{\text{SUSY}}}{M_W}$

Corrections to M_W , $\sin^2 \theta_{\text{eff}}$ → approximation via the ρ -parameter:

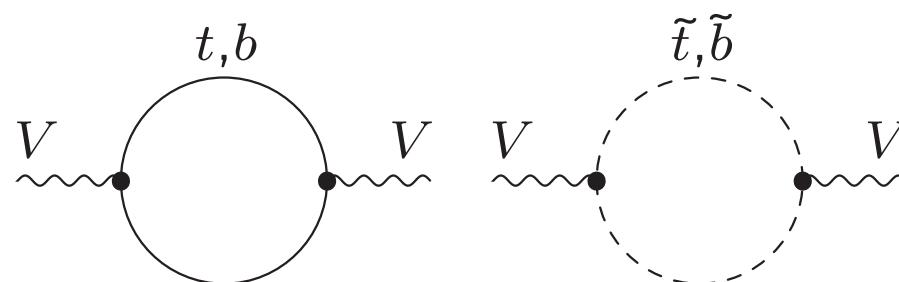
ρ measures the relative strength between
neutral current interaction and charged current interaction

$$\rho = \frac{1}{1 - \Delta\rho} \quad \Delta\rho = \frac{\Sigma_Z(0)}{M_Z^2} - \frac{\Sigma_W(0)}{M_W^2}$$

(leading, process independent terms)

$\Delta\rho$ gives the main contribution to EW observables:

$$\Delta M_W \approx \frac{M_W}{2} \frac{c_W^2}{c_W^2 - s_W^2} \Delta\rho, \quad \Delta \sin^2 \theta_W^{\text{eff}} \approx - \frac{c_W^2 s_W^2}{c_W^2 - s_W^2} \Delta\rho$$



$\Delta\rho^{\text{SUSY}}$ from \tilde{t}/\tilde{b} loops $> 0 \Rightarrow M_W^{\text{SUSY}} \gtrsim M_W^{\text{SM}}, \sin^2 \theta_{\text{eff}}^{\text{SUSY}} \lesssim \sin^2 \theta_{\text{eff}}^{\text{SM}}$

$$\Delta\rho^{\text{SUSY}} \text{ from } \tilde{t}/\tilde{b} \text{ loops} > 0 \quad \Rightarrow M_W^{\text{SUSY}} \gtrsim M_W^{\text{SM}}, \sin^2 \theta_{\text{eff}}^{\text{SUSY}} \lesssim \sin^2 \theta_{\text{eff}}^{\text{SM}}$$

SM result for M_W , $\sin^2 \theta_{\text{eff}}$:

- full one-loop
- full two-loop
- leading 3-loop via $\Delta\rho$ (not yet n_f^3 ;-)
- leading 4-loop via $\Delta\rho$

Best MSSM result for M_W :

[S.H., W. Hollik, G. Weiglein, L. Zeune '13]

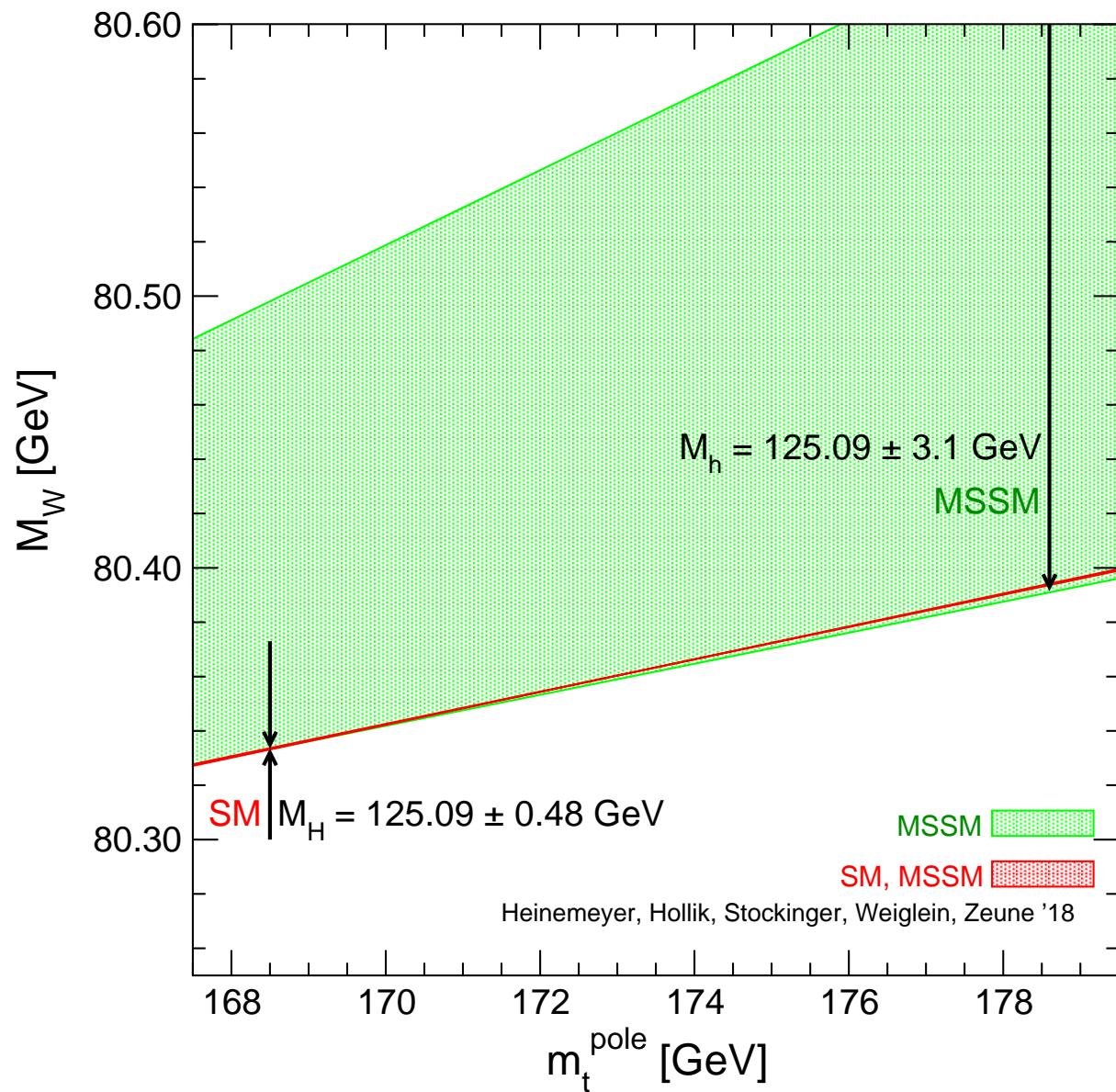
- full SM result (via fit formel)
- full MSSM one-loop (incl. CPV phases)
- all existing two-loop $\Delta\rho$ contributions ($\Delta\rho^{\alpha\alpha_s}$, $\Delta\rho^{\alpha_t^2, \alpha_t\alpha_b, \alpha_b^2}$)

$\Delta\rho$ does not contain:

- effects from sleptons
- effects from charginos/neutralinos

⇒ non- $\Delta\rho$ one-loop and $\Delta\rho$ two-loop contributions
sometimes non-negligible!

Example: Prediction for M_W in the SM and the MSSM :
[S.H., W. Hollik, D. Stockinger, G. Weiglein, L. Zeune '18]



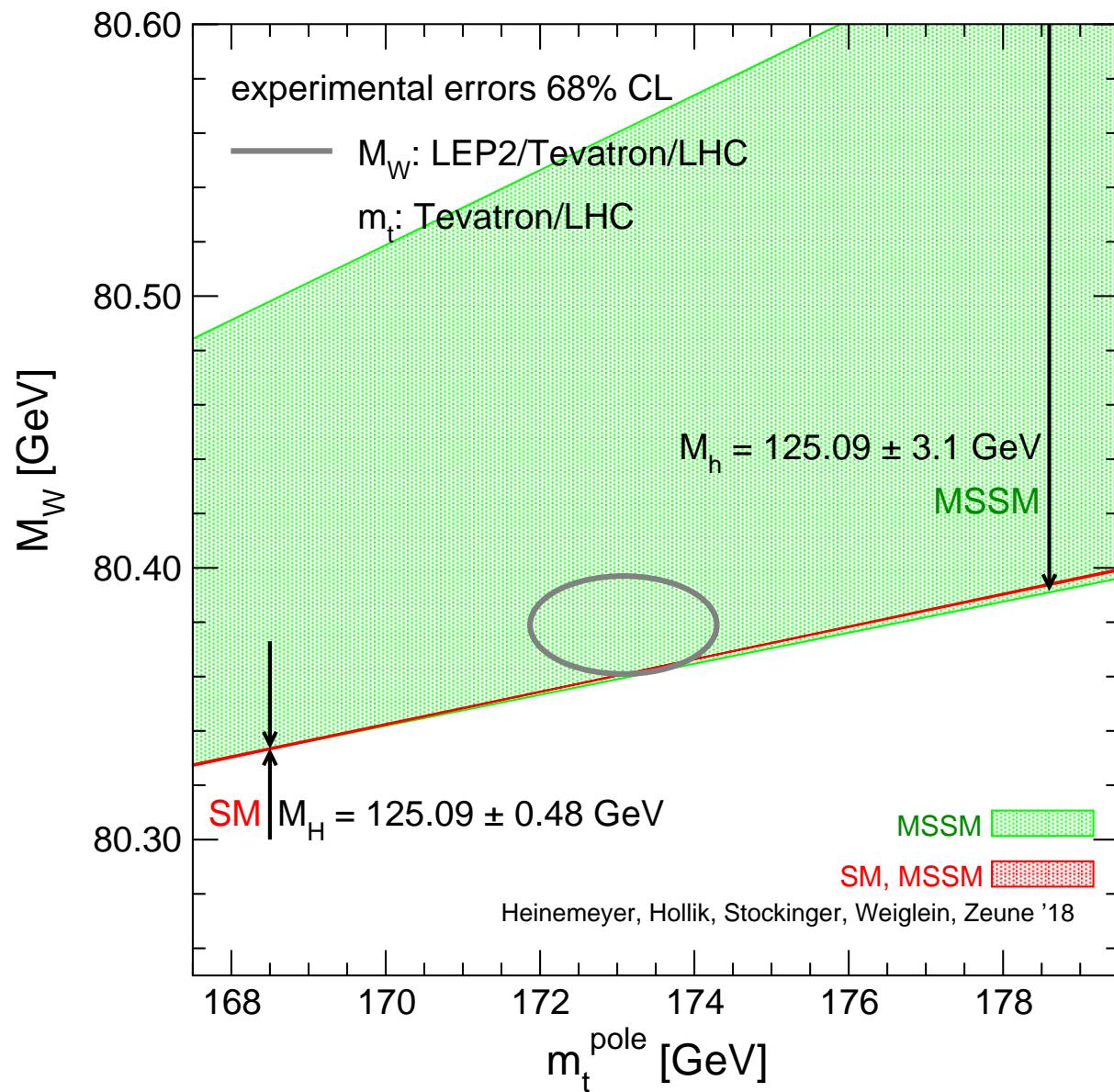
MSSM band:
scan over
SUSY masses

overlap:
SM is MSSM-like
MSSM is SM-like

SM band:
variation of M_H^{SM}

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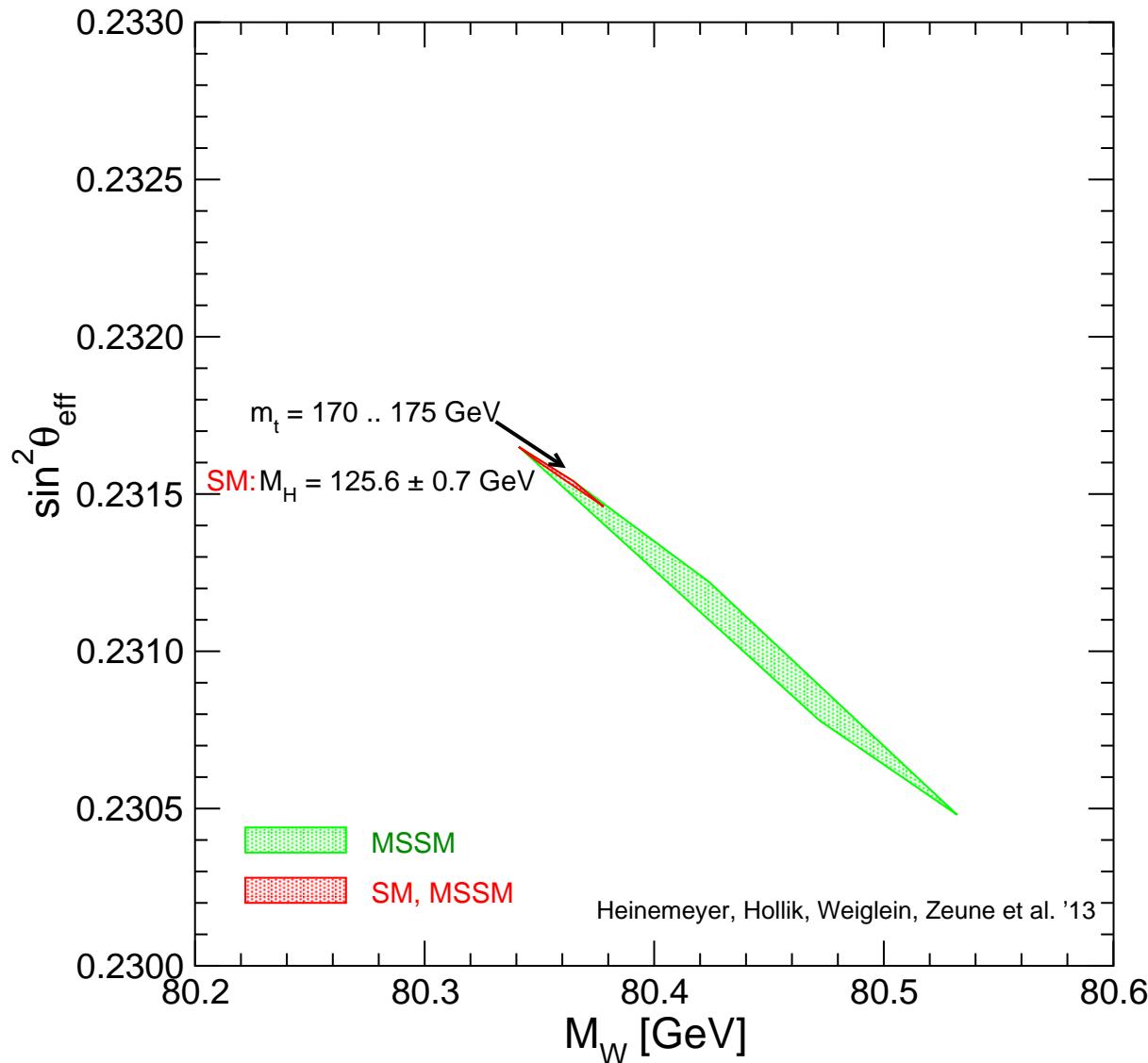
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Example: Prediction for M_W and $\sin^2 \theta_{\text{eff}}$ in the **SM** and the **MSSM** :
[S.H., W. Hollik, G. Weiglein, L. Zeune et al. '13]

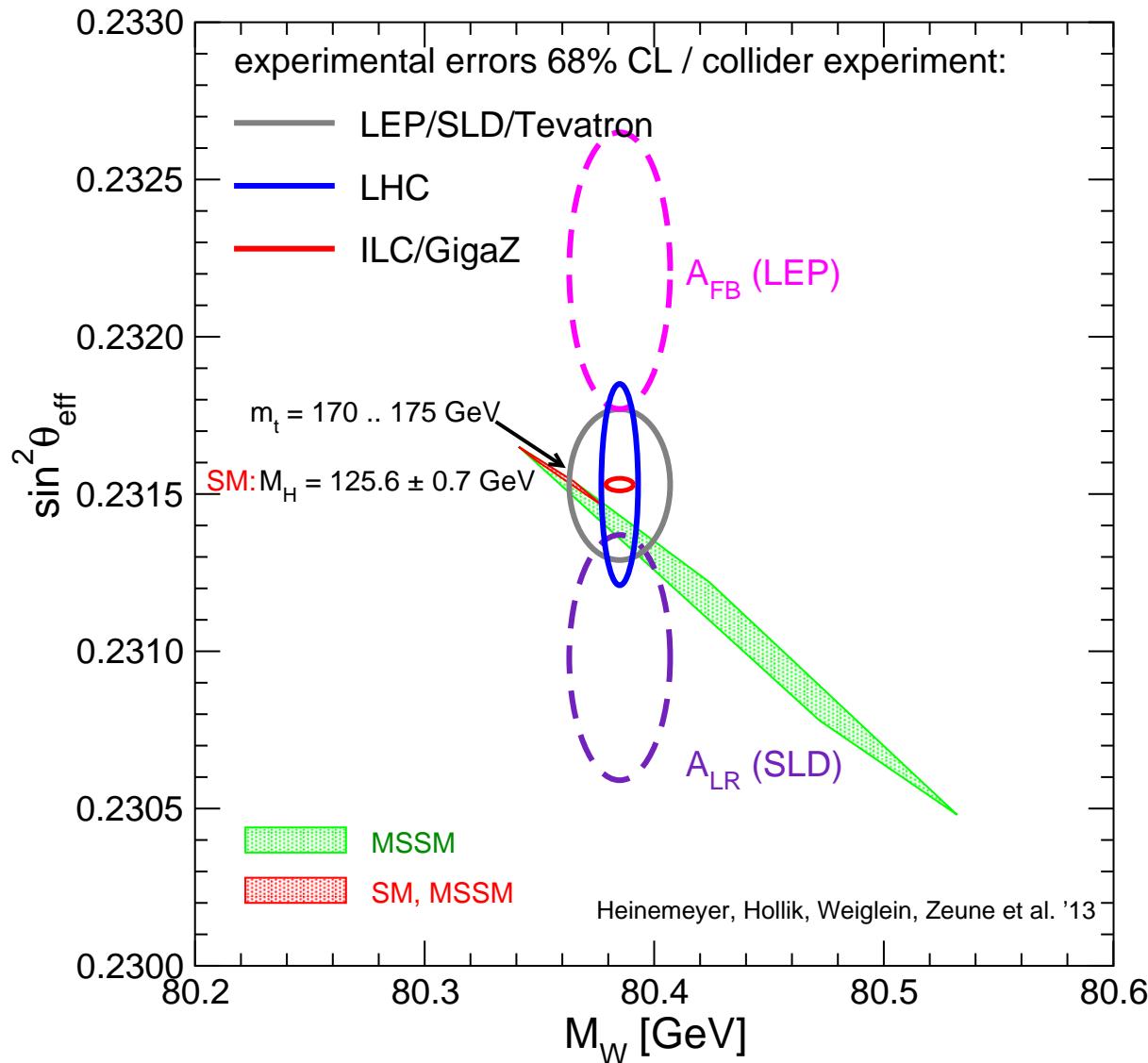


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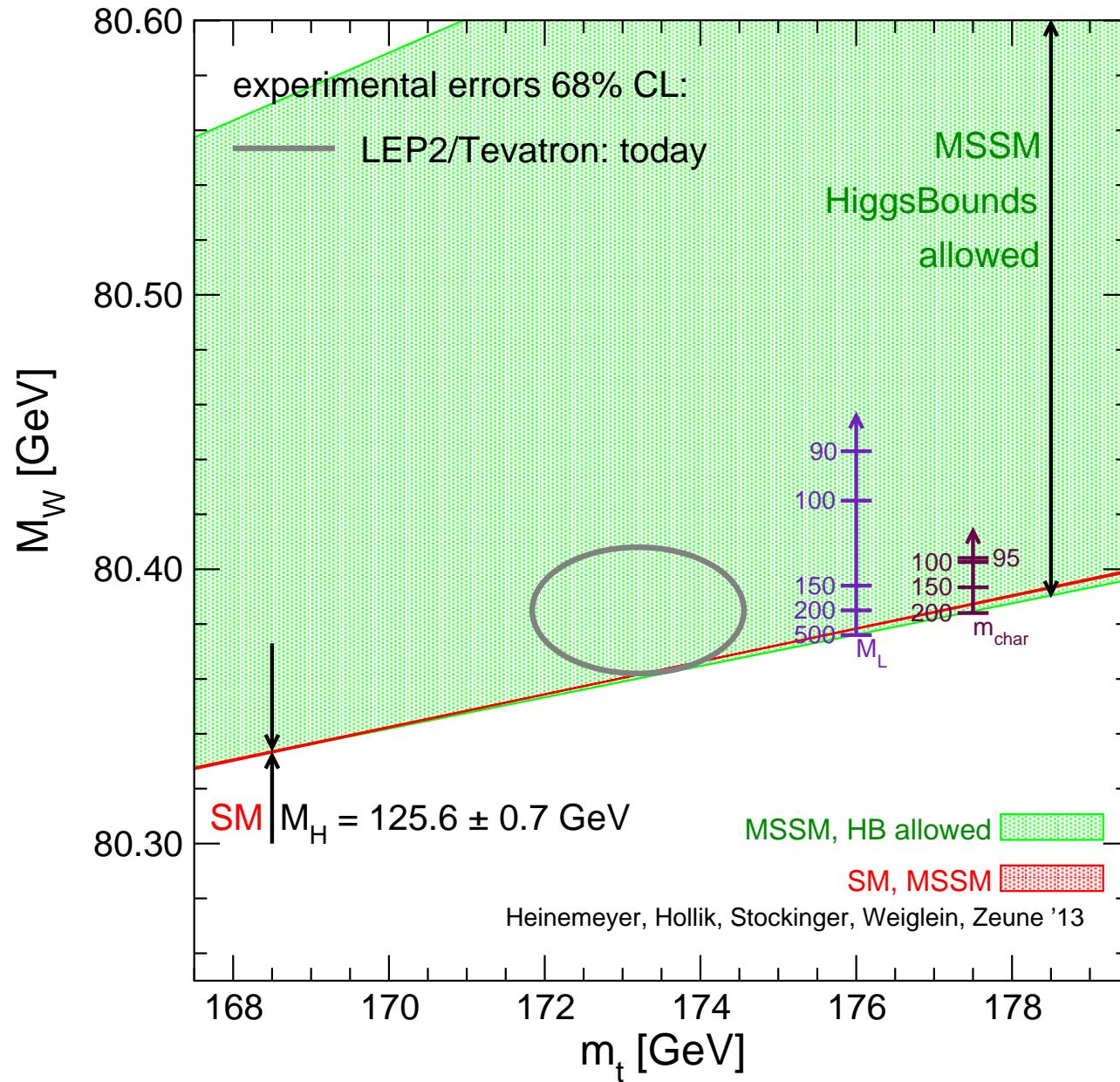
Example MSSM scenario (I):

[S.H., G. Weiglein, L. Zeune '13]

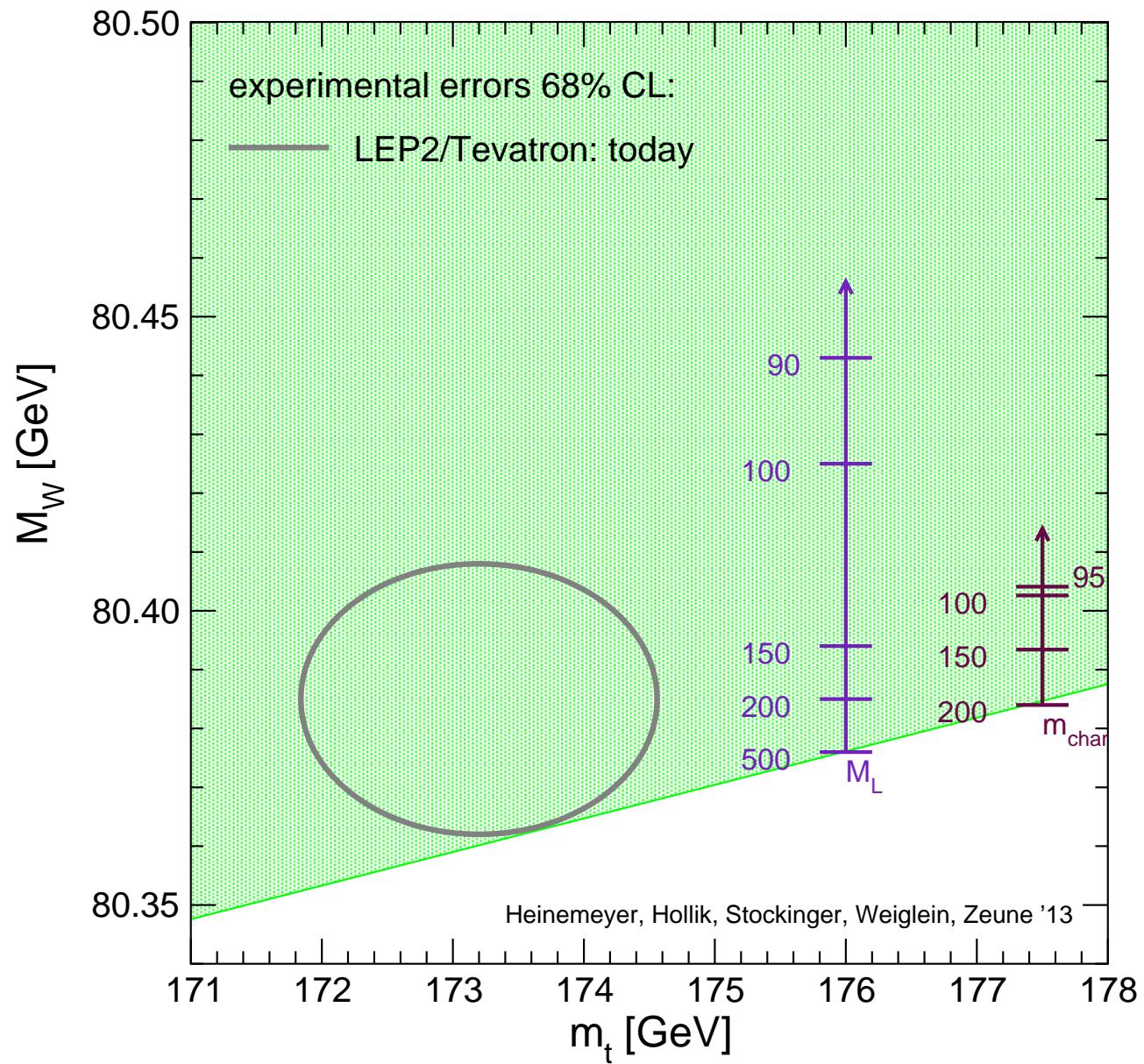
⇒ extensive parameter scan:

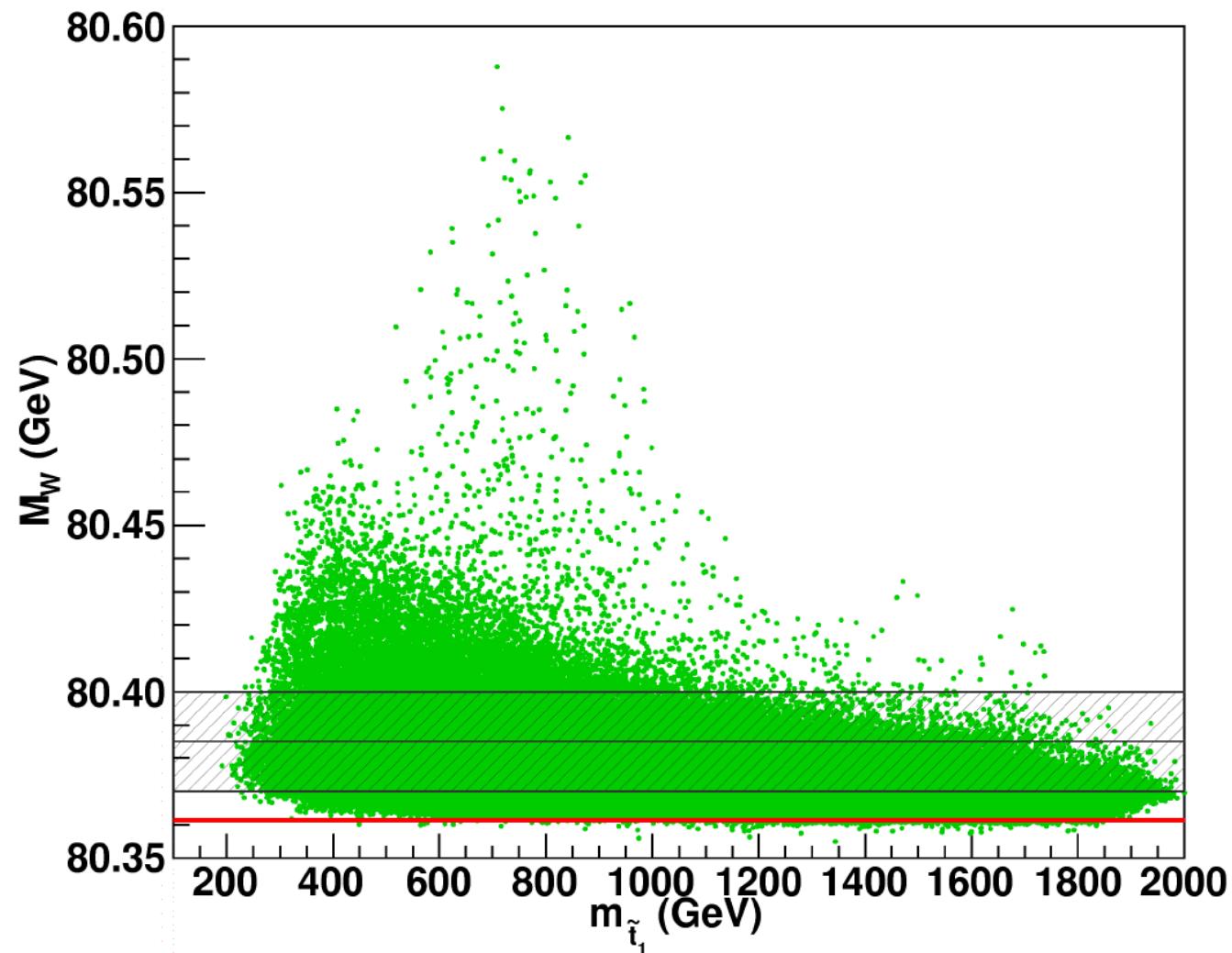
Parameter	Minimum	Maximum
μ	-2000	2000
$M_{\tilde{E}_{1,2,3}} = M_{\tilde{L}_{1,2,3}}$	100	2000
$M_{\tilde{Q}_{1,2}} = M_{\tilde{U}_{1,2}} = M_{\tilde{D}_{1,2}}$	500	2000
$M_{\tilde{Q}_3}$	100	2000
$M_{\tilde{U}_3}$	100	2000
$M_{\tilde{D}_3}$	100	2000
$A_e = A_\mu = A_\tau$	$-3 M_{\tilde{E}}$	$3 M_{\tilde{E}}$
$A_u = A_d = A_c = A_s$	$-3 M_{\tilde{Q}_{12}}$	$3 M_{\tilde{Q}_{12}}$
A_b	$-3 \max(M_{\tilde{Q}_3}, M_{\tilde{D}_3})$	$3 \max(M_{\tilde{Q}_3}, M_{\tilde{D}_3})$
A_t	$-3 \max(M_{\tilde{Q}_3}, M_{\tilde{U}_3})$	$3 \max(M_{\tilde{Q}_3}, M_{\tilde{U}_3})$
$\tan \beta$	1	60
M_3	500	2000
M_A	90	1000
M_2	100	1000

Example MSSM scenario (I): effects beyond $\Delta\rho$: EW particles

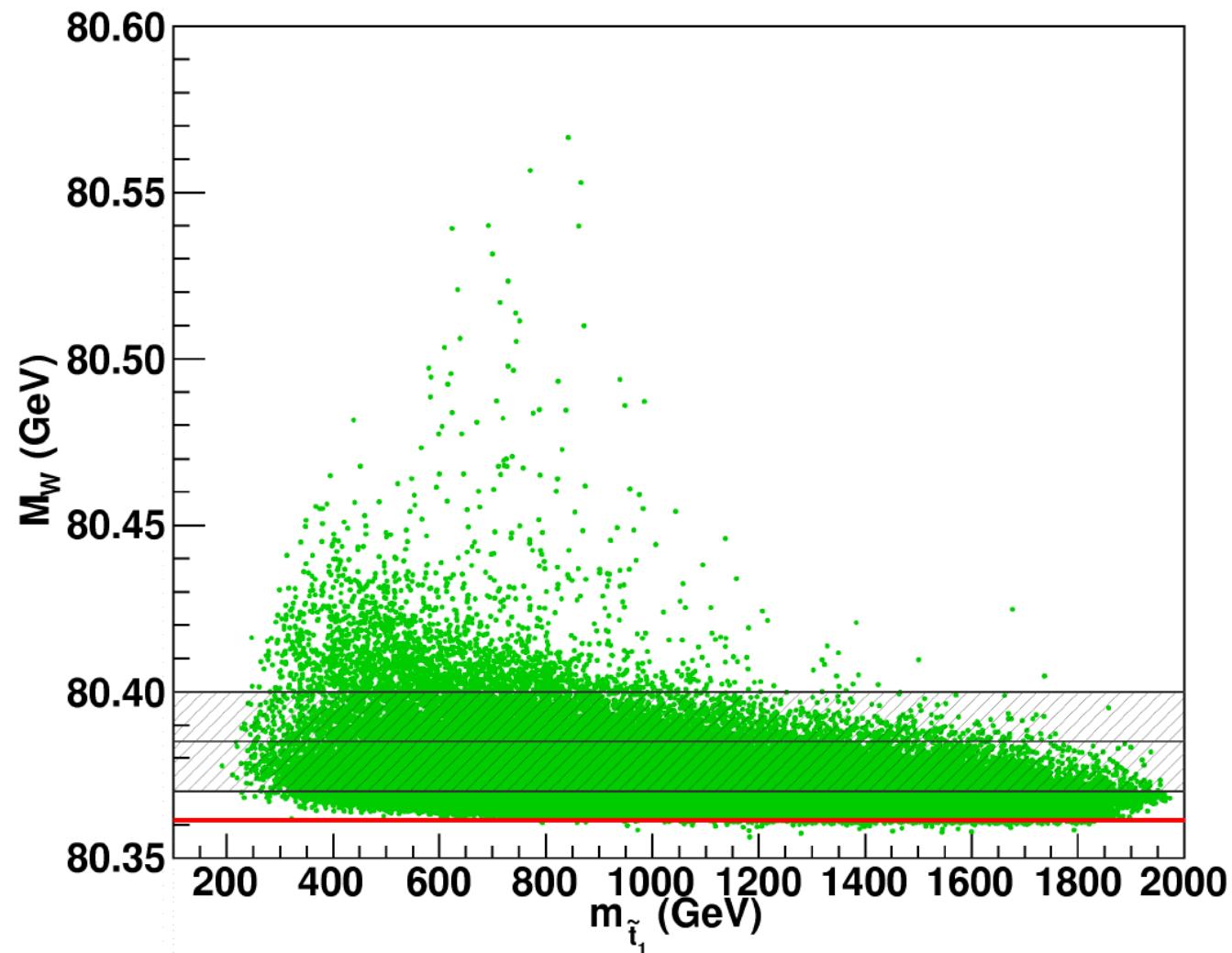


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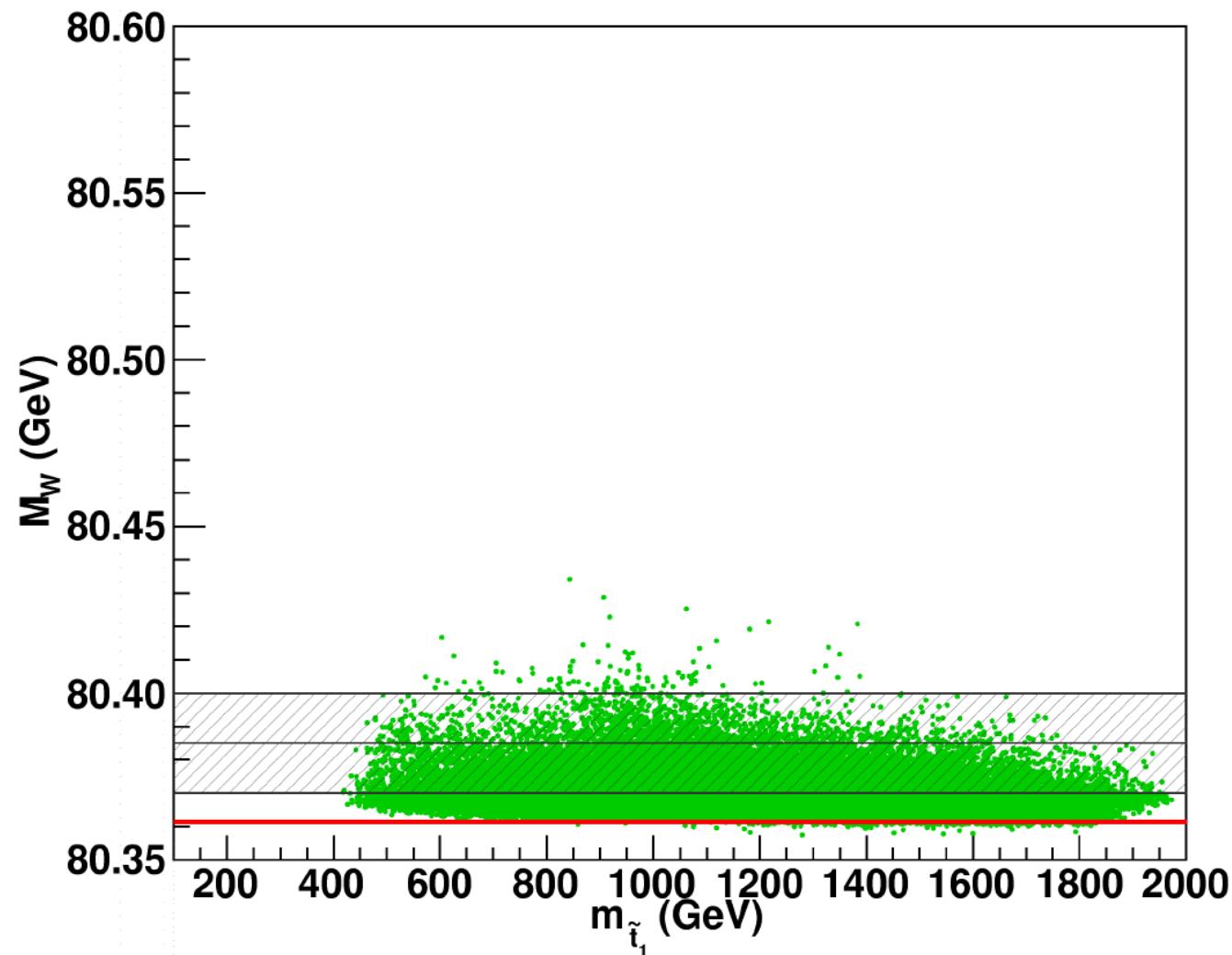




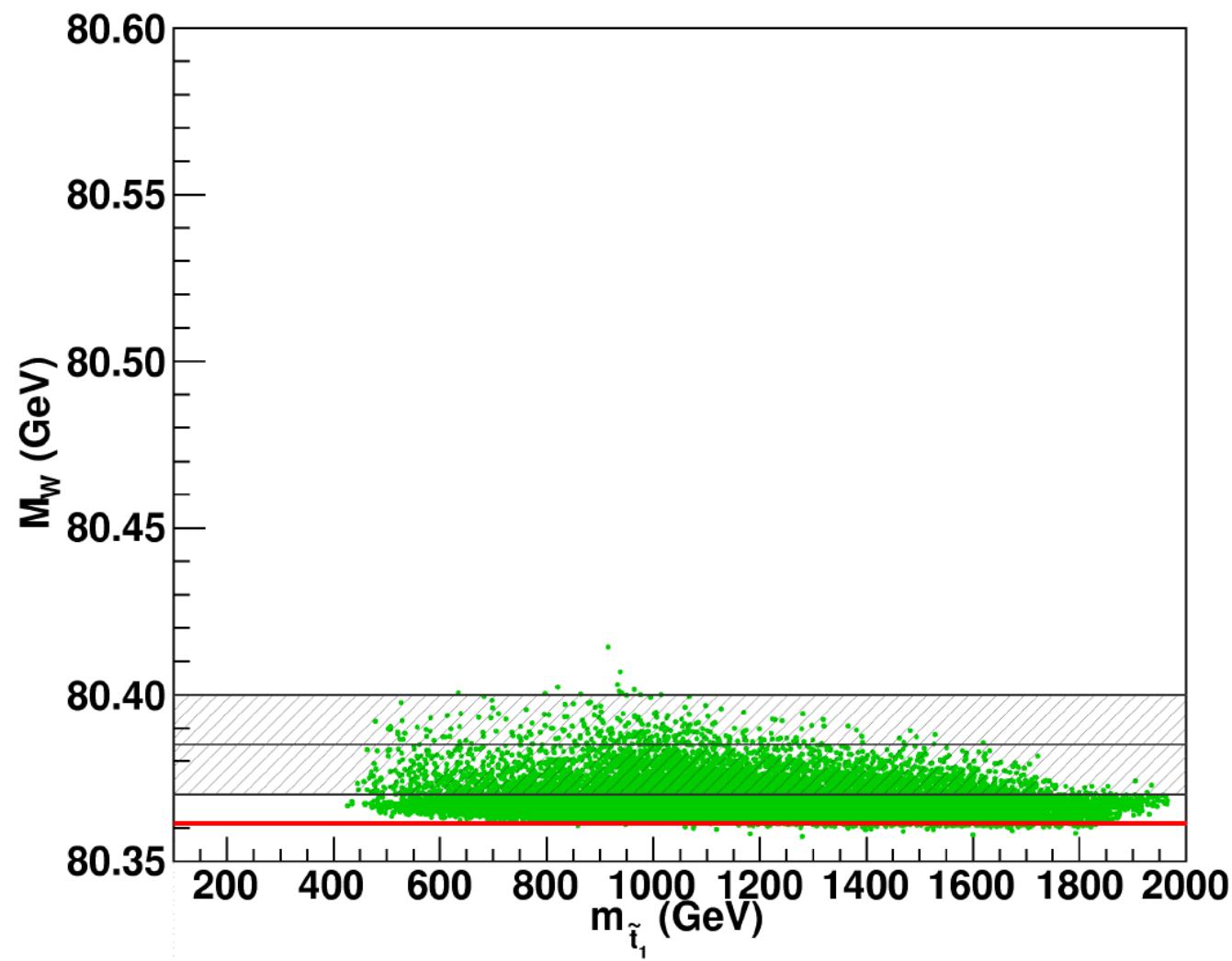
All points HiggsBounds allowed



$\dots \oplus m_{\tilde{q}_{1,2}}, m_{\tilde{g}} > 1200 \text{ GeV}$



$\dots \oplus m_{\tilde{b}_i} > 500 \text{ GeV}$



$\dots \oplus m_{\tilde{l}}, m_{\tilde{\chi}_i^\pm}, m_{\tilde{\chi}_j^0} > 500 \text{ GeV}$

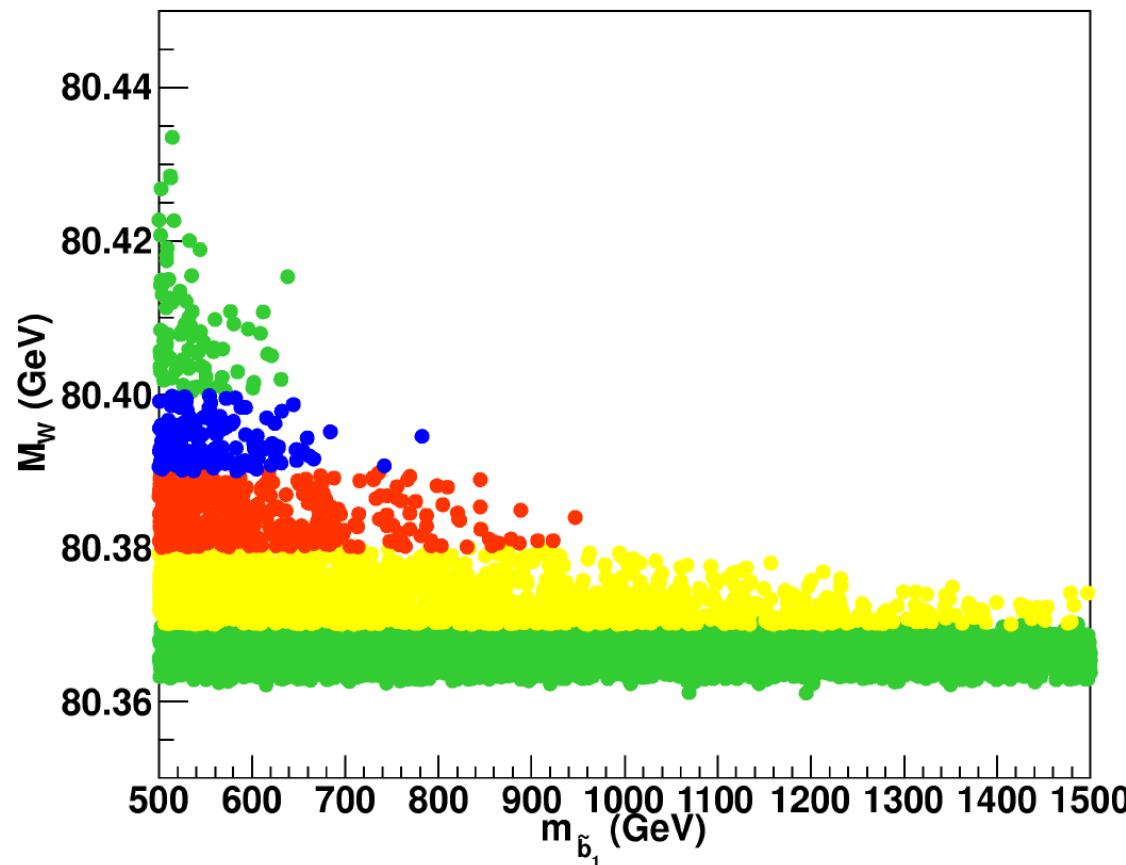
\Rightarrow update worthwhile?!

Example MSSM scenario (II):

[S.H., G. Weiglein, L. Zeune '15]

$m_{\tilde{t}_1} = 400 \pm 40$ GeV, Other masses $\gtrsim 500$ GeV

$M_W^{\text{exp}} = 80.375 \pm 0.005$ GeV, 80.385 ± 0.005 GeV, 80.395 ± 0.005 GeV



⇒ precision below 5 MeV required!

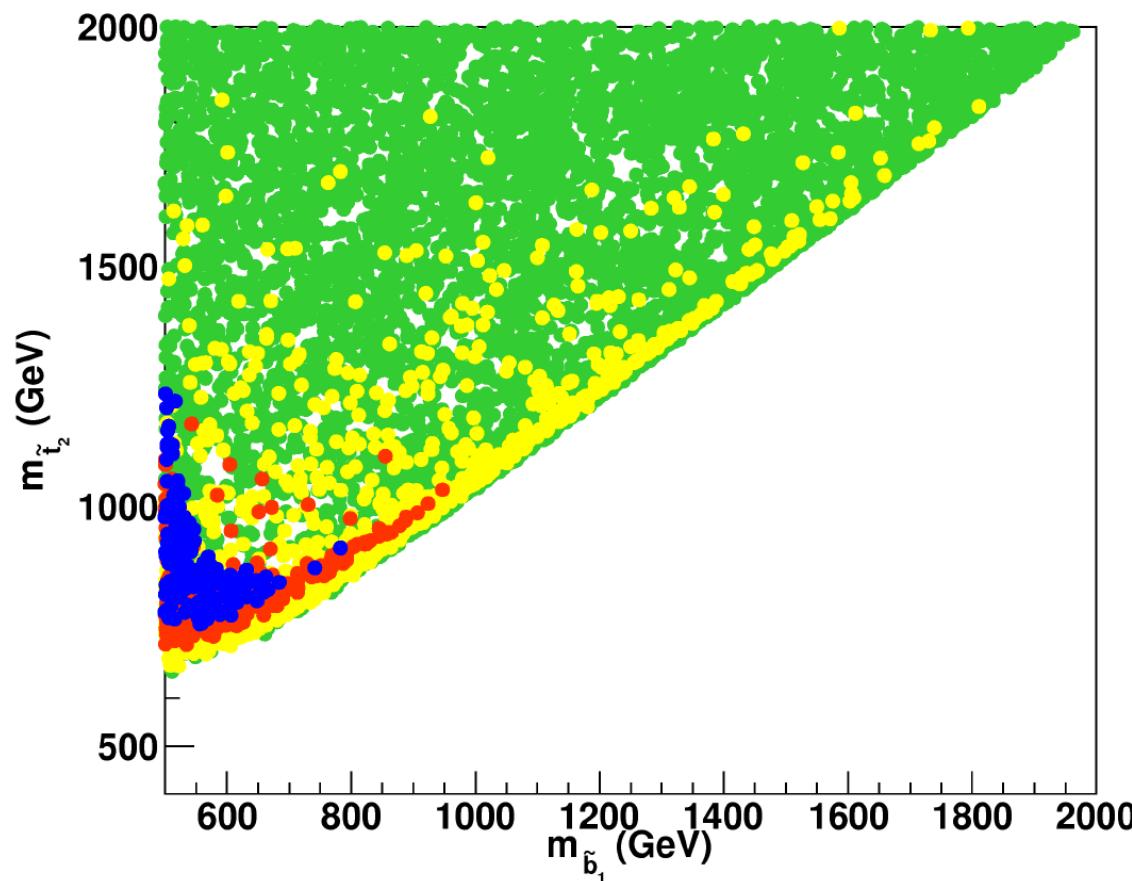
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3. Why not (only) EFTs?

EFTs have many virtues!

But they also have many (un?)known short-comings:

- all BSM physics is assumed to be very heavy,
out of the direct reach of current/future colliders
(Not my favorite future physics scenario . . .)
- if one finds large effects in the EFT predictions,
it is not clear whether this can be reproduced by any real model
- EFTs as such leave unclear to what underlying real model
a certain effect corresponds
⇒ this requires a “model by model” prediction for the EFT
- . . .

Let us assume that we do see a deviation from the SM

What do we learn from that?

How do we learn something from that?

- ⇒ We have to compare the **observed** deviation with
predicted deviations
- ⇒ Preferably with the predicted deviations in a **concrete models**

We want to learn which physics is responsible for the deviations

- ⇒ A comparison with an **EFT result** subsequently requires the mapping to **concrete models** anyway!

Needed:

- sufficiently precise predictions in **BSM** model
- ... including uncertainty estimates
- Analysis of patterns of deviations?!

4. Conclusions

- EWPOs are a powerful tool to learn about unknown physics scales
⇒ top mass and Higgs mass were predicted correctly (within the SM)
- “Classical” EWPO: M_W , $\sin^2 \theta_{\text{eff}}$, Γ_Z , R_l , R_q , . . .
Do not forget possible new EWPOs: a_μ , M_h , . . .
- Current predictions within the SM are sufficiently under control for now
- Situation is less clear for BSM models
⇒ better overview necessary! ⇒ Snowmass 2021?
⇒ Analysis of patterns of deviations?!
- MSSM as a showcase:
 - sufficiently worked out, including rough uncertainty estimate
 - possible clear patterns
 - $\Delta\rho$ not sufficient to cover the possible effects
 - EWPOs can tell us about other unknown scales of the MSSM
- Assume an observed deviation from the SM:
 - compare observed deviations with predicted deviations
 - preferably in a concrete model
 - Comp. with EFT requires subsequent mapping to concrete models



Further Questions?